

## Comments to the diagrams

The diagram of "reception performance" shows the signal-to-noise ratio of extra terrestrial signal reception (sun noise) dependent on f/D ratio.

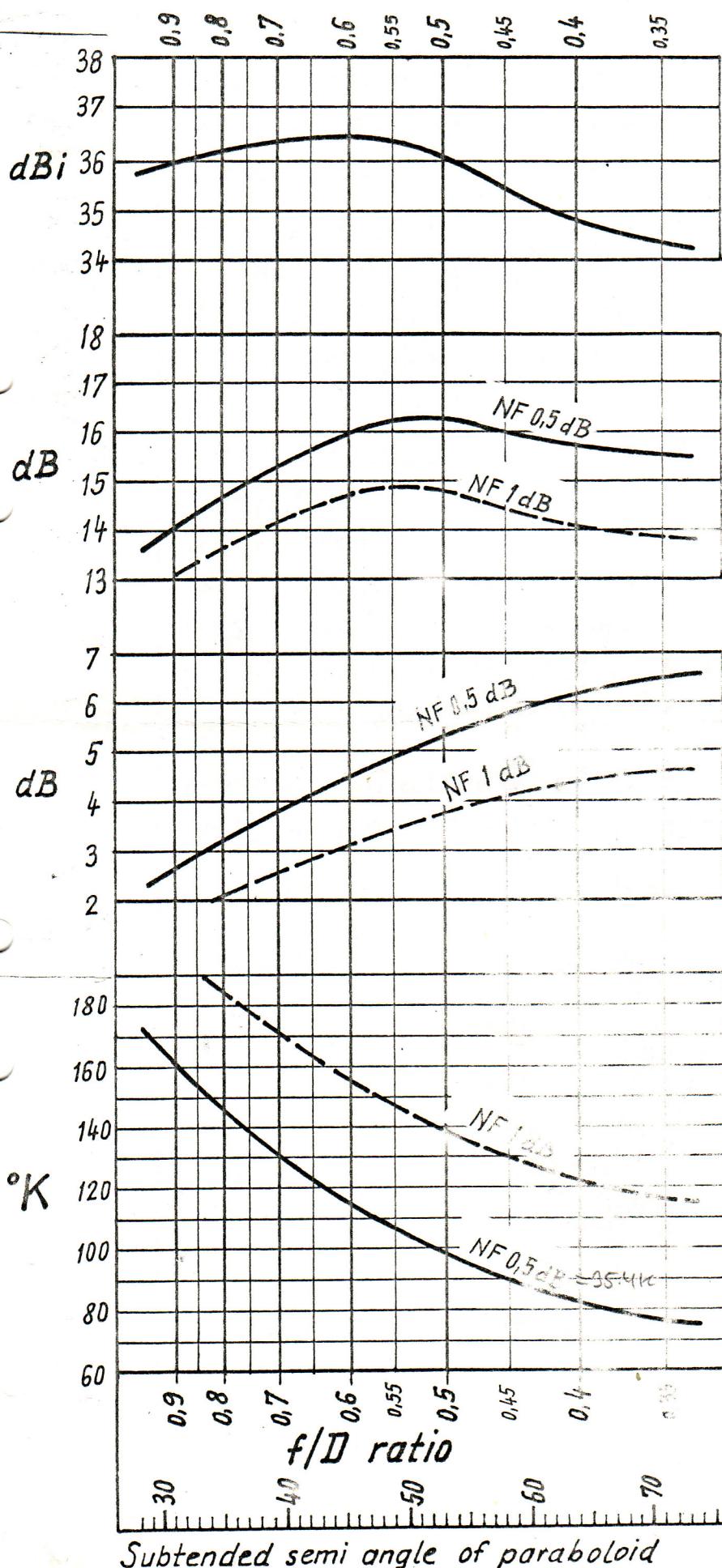
Using a 0,5 dB NF front end, the best performance of the receiving system looks at 0,52 f/D (corresponding to -12,6 dB dish edge illumination). The maximum gain is achieved at 0,59 f/D (corresponds to -9,9 dB dish edge illumination).

The best overall performance requires a compromise. In this case, compromise means an f/D ratio which gives equalized loss of transmission- and reception-performance. A closer look at the diagram shows this compromise at 0,55 f/D (both losses approx 0,2 dB) which corresponds to 11,3 dB aperture taper.

However, small movements in the 0,5 - 0,6 f/D range achieves no significant effects on performance.

June 84, OE 9 PM J

Performance of a dish antenna in relation to f/D ratio,  
illuminated by a dual mode horn (W2IMU design) on 1296 Mc



TRANSMISSION PERFORMANCE

Gain of 2oft dish dia

On 0,59 f/D maximum aperture efficiency  
 $\eta_f = 0,645$ , where spill-over  $\eta^s = 0,831$  and illumination  $\eta^i = 0,776$

RECEPTION PERFORMANCE

maximum at 0,52 (0,54)  
f/D ratio

Calculated sun noise (Flux 65) for a 2oft dish and 0.5(1.0) dB noise figure preamp.

Ratio of thermal earth noise (antenna aimed at zenith) to sky reception.

It is a measure of receiver system temperature independent of antenna gain.

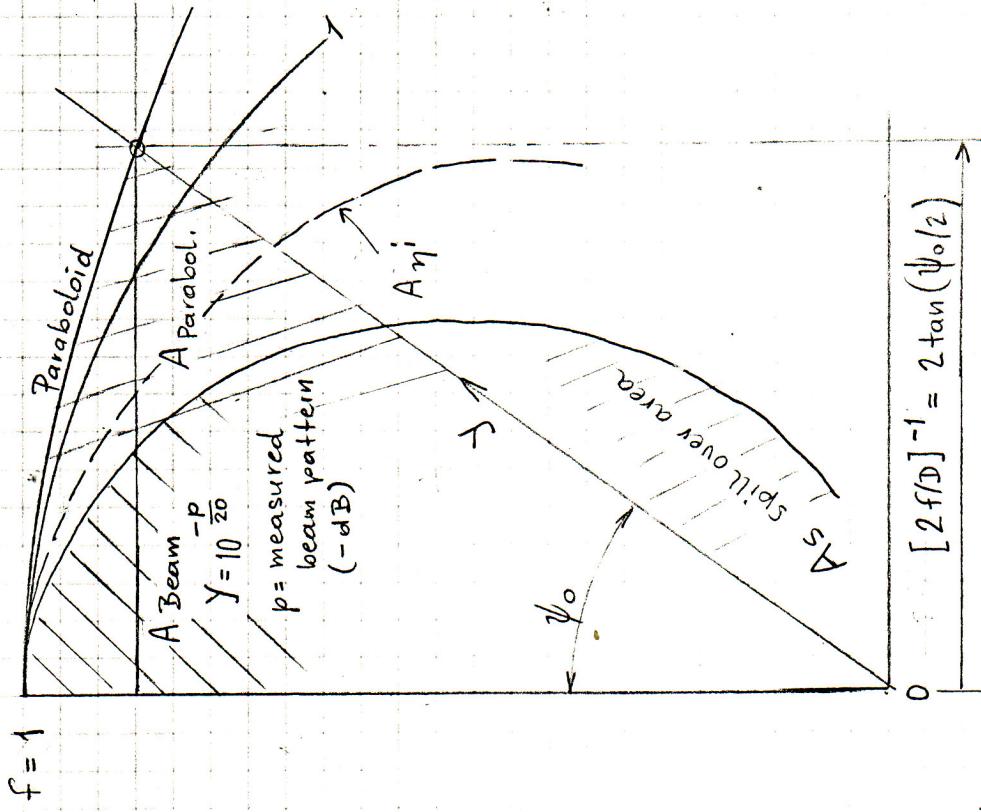
Noise temperature of receiving system with 0.5(1.0) dB NF front end, antenna aimed at zenith.

All values are calculated based on a measured beam pattern diagram of the feed horn.

Reflector-, phase- and polarization efficiencies are  $\eta = 1$ .

## SCHEME FOR CALCULATION

of illumination efficiency  $\eta^i$  and spill over efficiency  $\eta^s$  of a paraboloid antenna illuminated by a radiator with equal E- and H-plane beam pattern (W21MU-horn)



$$\begin{aligned}
 A_P &= (4f/D)^{-1} + [192(f/D^3)]^{-1} && * \text{ Simpson's approximation} \\
 &= \tan(\psi_0/2) + \frac{[\tan(\psi_0/2)]^3}{3} \\
 \frac{h}{3} &= \frac{\pi \cdot \psi_0}{3 \cdot 360 \cdot n}; n = 2k
 \end{aligned}$$
  

$$\begin{aligned}
 A_B &= \int_0^{\psi^0} f(\psi) d\psi \approx \frac{h}{3} (\gamma_0 + 4\gamma_1 + 2\gamma_2 + 4\gamma_3 + 2\gamma_4 + \dots) \\
 &\quad + 2\gamma_{2k-2} + 4\gamma_{2k-1} + \gamma_{2k} \\
 A_{\eta^i} &= \sqrt[3]{A_B \cdot A_P^2}; \quad \eta^i = \frac{A_{\eta^i}}{A_P} = \sqrt[15]{\frac{A_B}{A_P}}
 \end{aligned}$$
  

$$A_s = \int_{\psi^0}^{180^\circ} f(\psi) d\psi; \quad \eta^s = \frac{A_B + A_s}{A_P}$$

Aperjur efficiency  $\eta^f = \eta^i \cdot \eta^s \cdot \eta^\phi \cdot \eta^R$ ;  $\eta^\phi$  (phase)  $\approx 1$ ,  $\eta^R$  (polarization)  $\approx 1$ ,  $\eta^R$  (reflector performance)  $\approx 0.85 \dots 0.995$

June 1984, OE9PHJ